

## KaVA (EAVN) astrometry for the Extreme Outer Galaxy Source G034.84-00.95

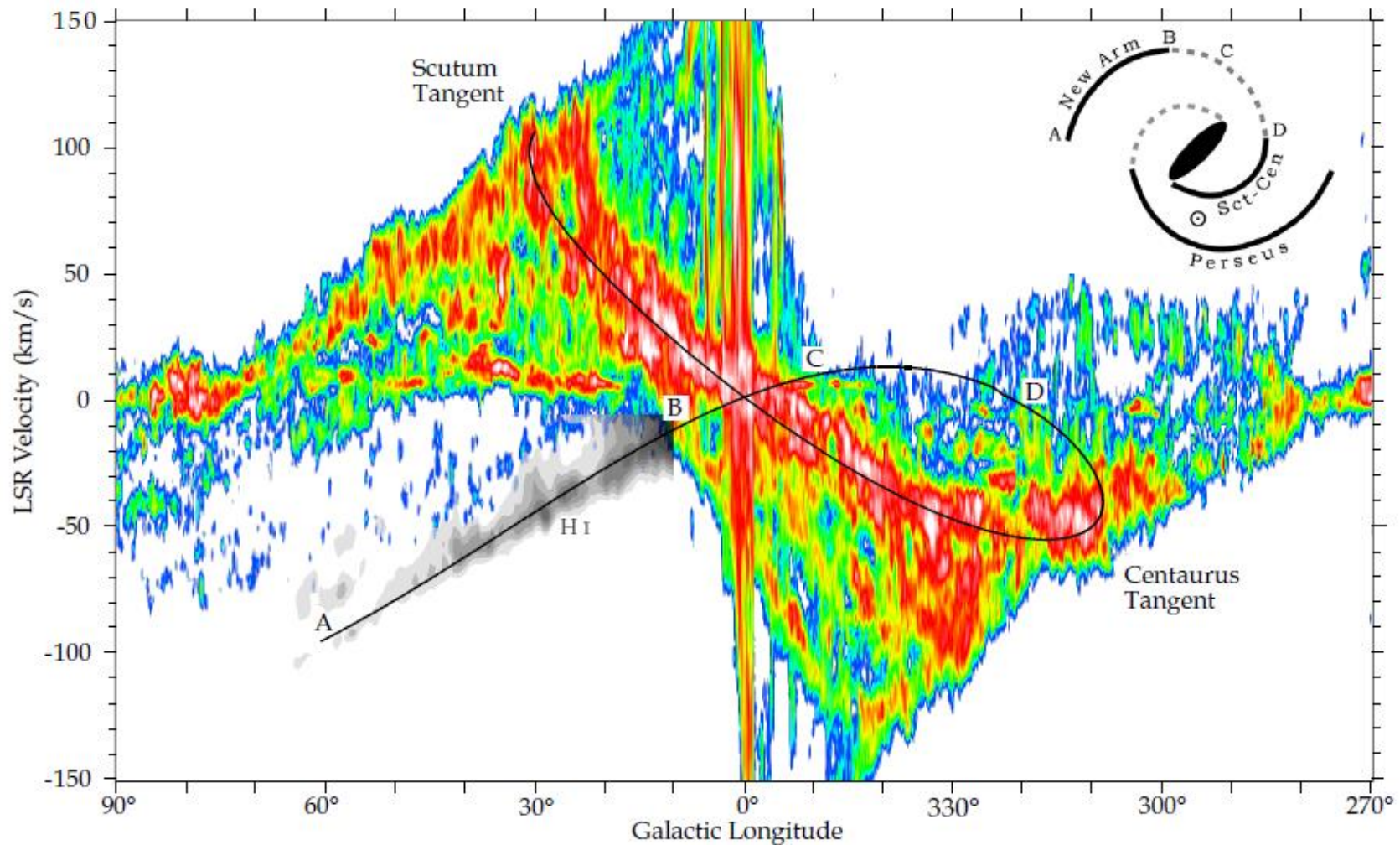
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N. Kawaguchi<sup>3</sup>, H. Imai<sup>4</sup>, J. Wu<sup>3</sup>, L. Cui<sup>5</sup>, S.-W. Kim<sup>1</sup>, P. Jiang<sup>5</sup>, et al.

September 24th, 2020@18<sup>th</sup> VERA UM, Zoom

<sup>1</sup> KASI; <sup>2</sup> SHAO; <sup>3</sup> NAOJ; <sup>4</sup> Kagoshima University; <sup>5</sup> XAO

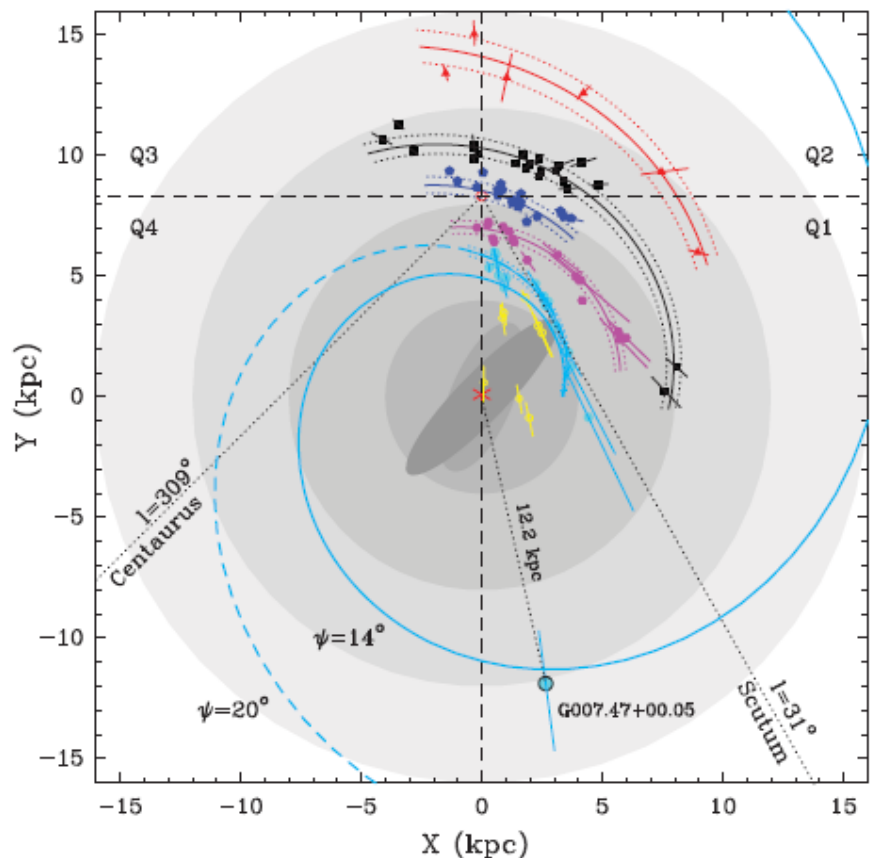
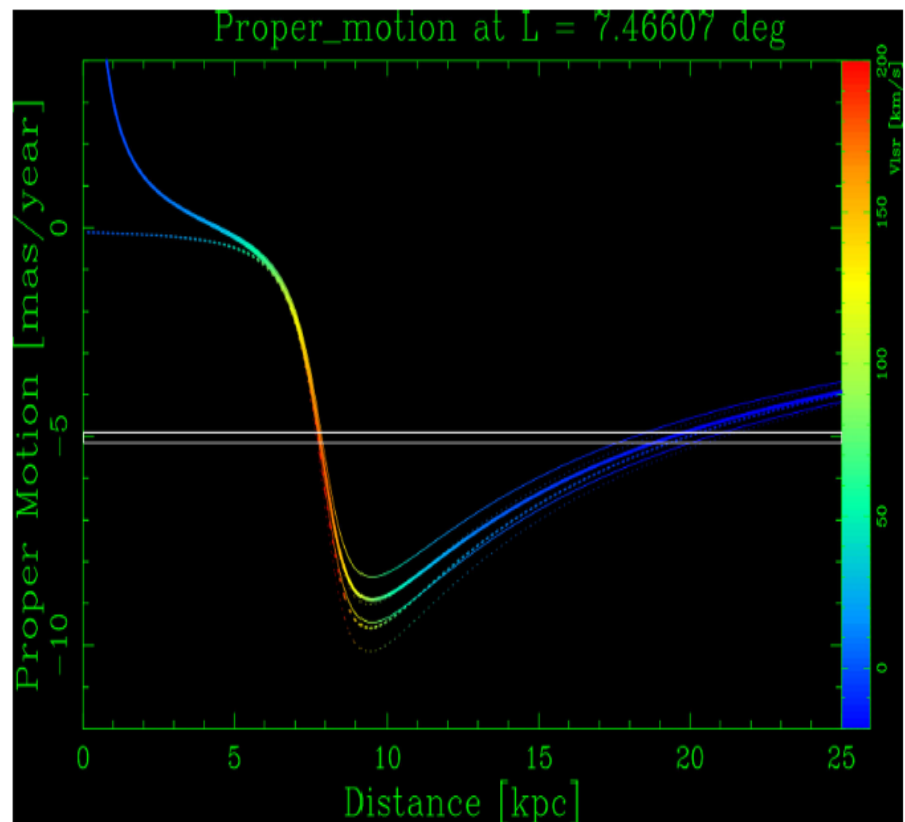


# A New Molecular Arm in the Extreme Outer Galaxy



- Lying ~15 kpc from the Galactic center traced by CO and H I
- The New arm is the continuous of the Scutum-Centaurus arm

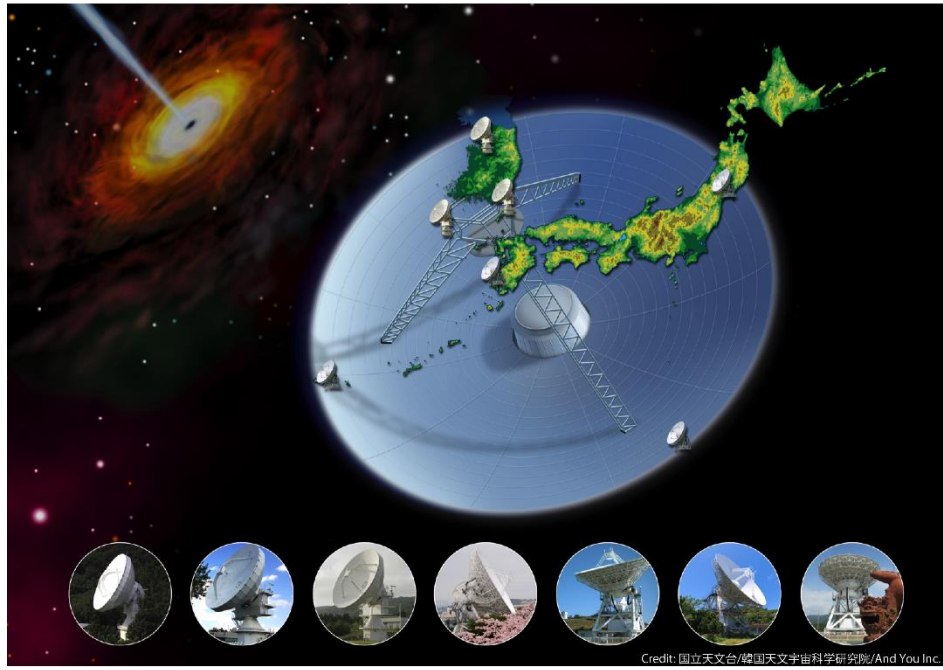
# VLBI Astrometry toward the Extreme Outer Galaxy



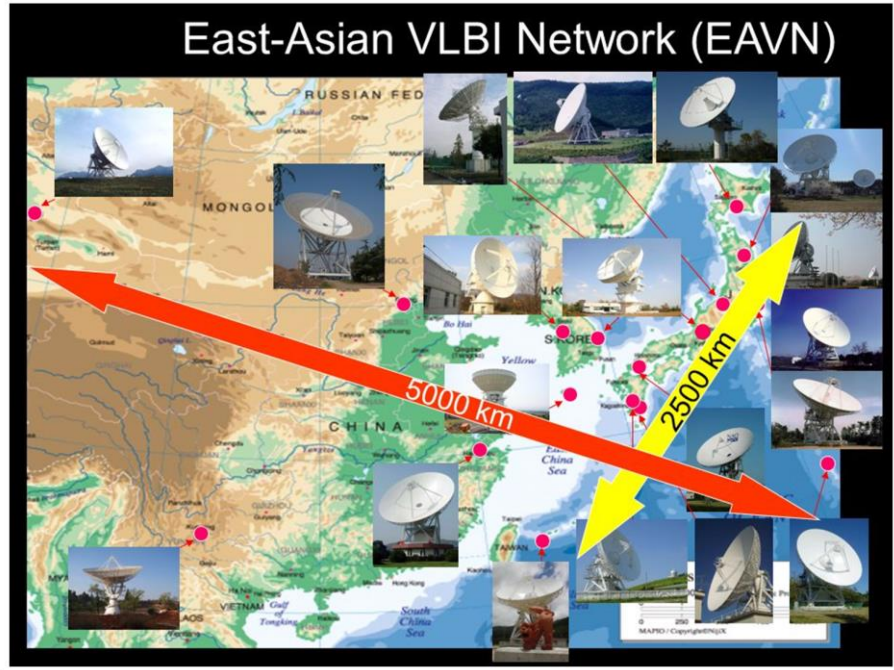
- Proper motion measurement with VERA
- $D = 20 \pm 2$  kpc for G7.47+0.06

- Parallax measurement with VLBA
- $D = 20.4 \pm 2.8 / -2.2$  kpc for G7.47+0.06

# Observations for the EOG source G034.84-00.95



Credit: 国立天文台/韓国天文宇宙科学研究所/And You Inc



Date: 2019/Sep ~ 2020/May (9 epochs)

VLBI array: KaVA

Frequency: 22 GHz (H<sub>2</sub>O maser)

Velocity spacing: 0.42 km s<sup>-1</sup>

Polarization: LHCP

Longest baseline: **2,300 km**

Mode: **Astrometry**

Date: 2020/Sep ~ 2020/Dec (approved)

VLBI array: EAVN

Frequency: 22 GHz (H<sub>2</sub>O maser)

Velocity spacing: 0.42 km s<sup>-1</sup>

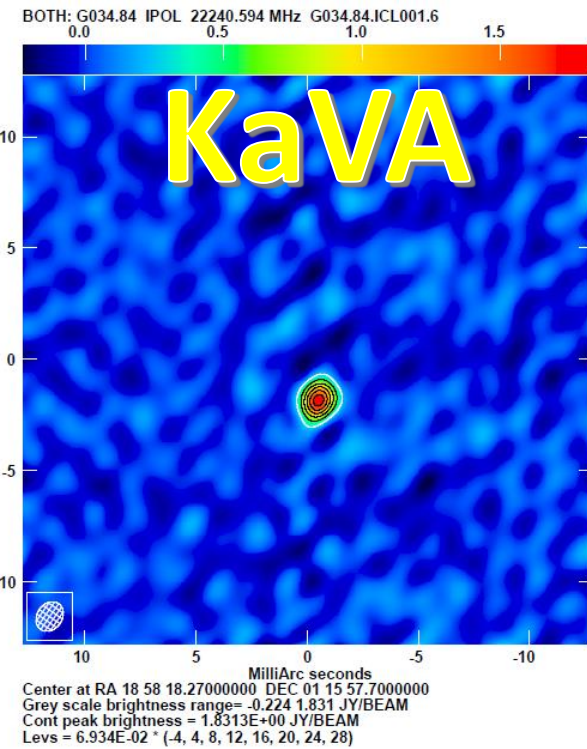
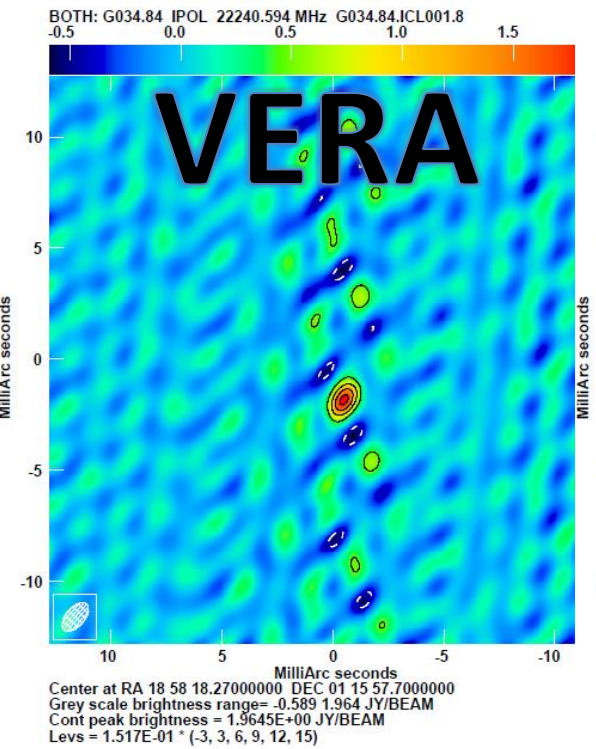
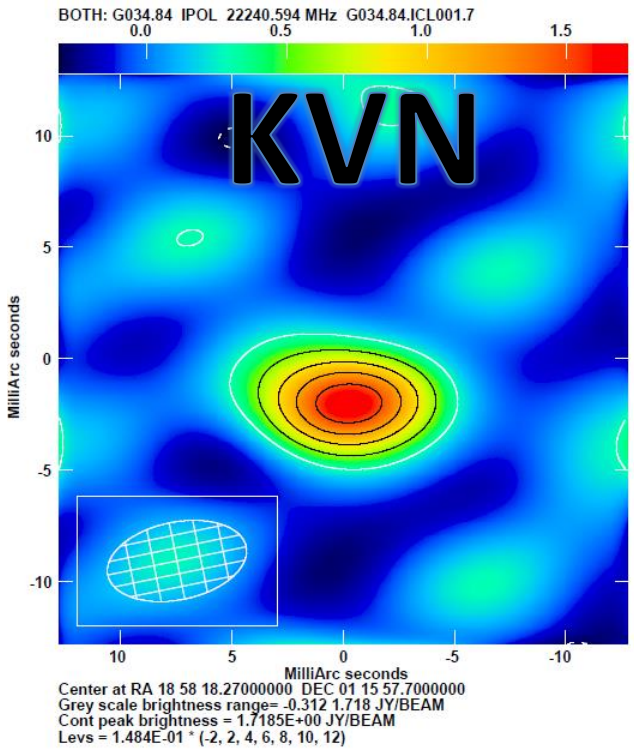
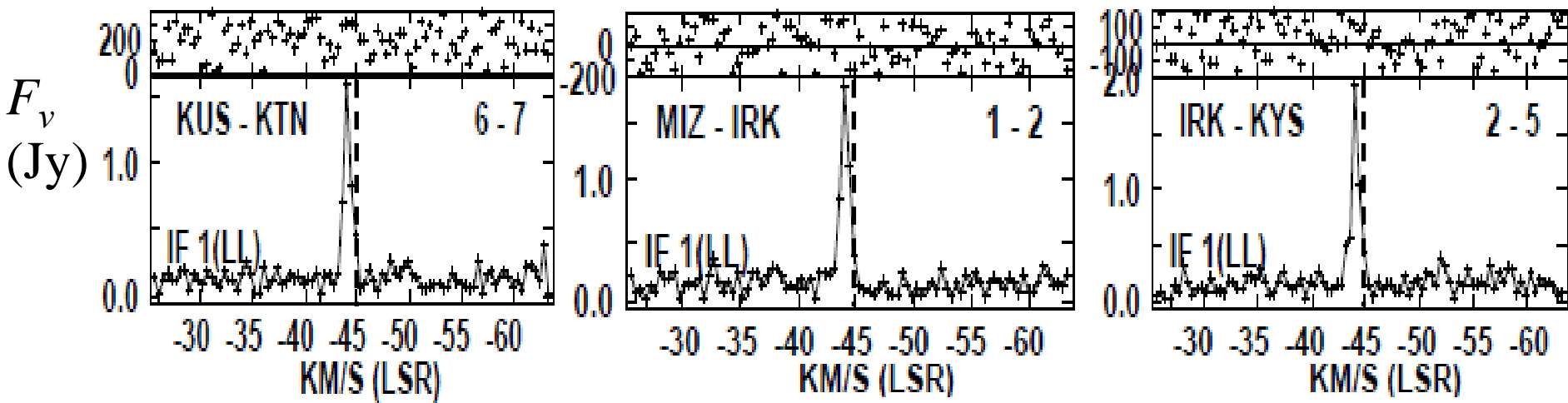
Polarization: LHCP

Longest baseline: **5,200 km**

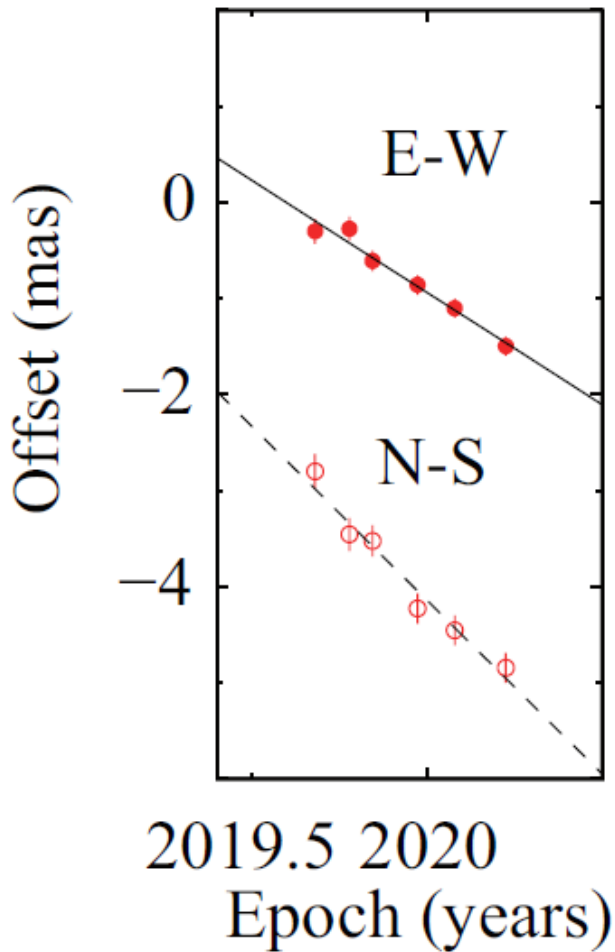
**(Ogasawara - Urumqi)**

Mode: **Astrometry**

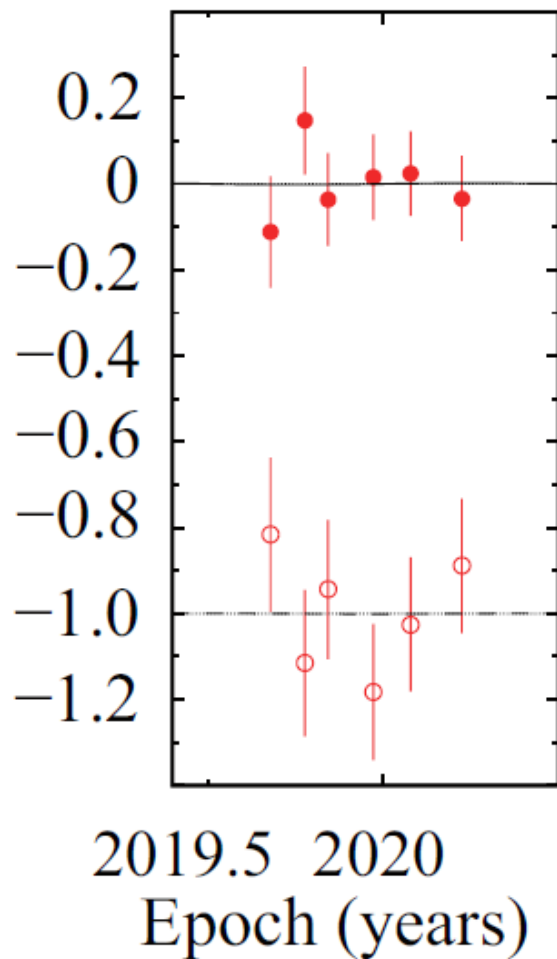
# Preliminary results with half year data



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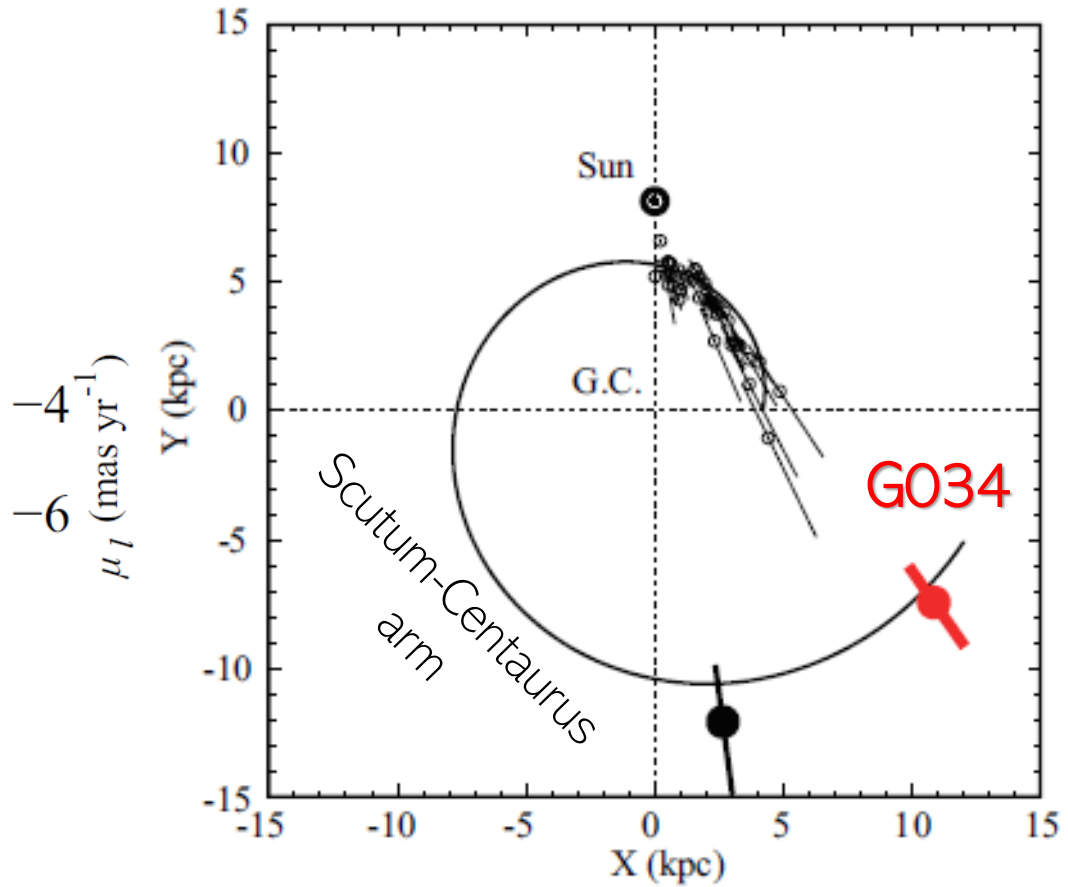
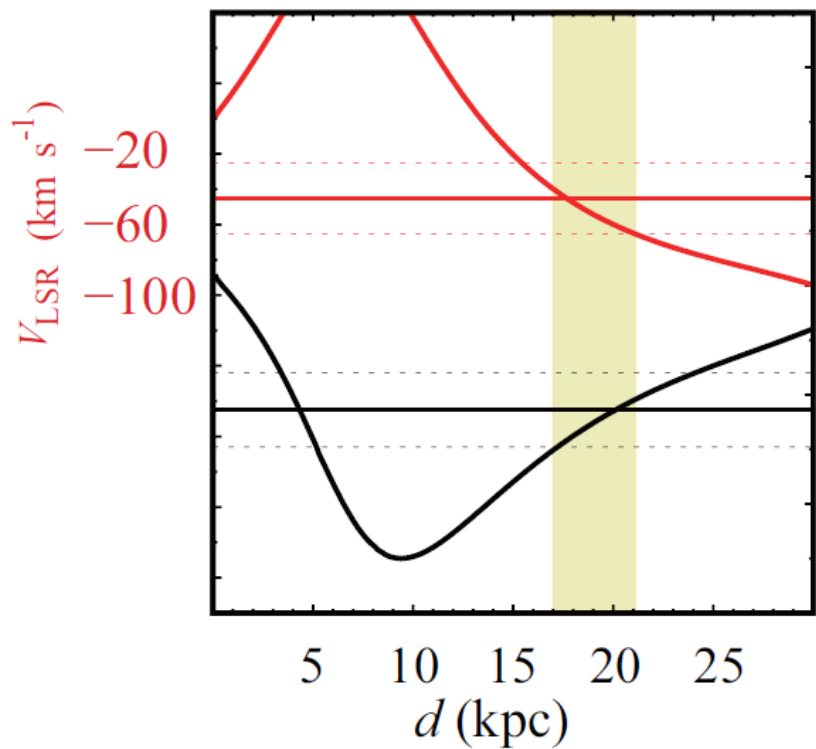


$\mu_\alpha \cos \delta = \blacksquare \text{ mas yr}^{-1}$   
 $\mu_\delta = \blacksquare \text{ mas yr}^{-1}$   
 \* Error decreases as  $t^{1.5}$



$\pi = \blacksquare \text{ mas}$   
 Insignificant parallax  
 due to the half year data ( $< 1 \text{ yr}$ ).

# Discussions with the Preliminary results



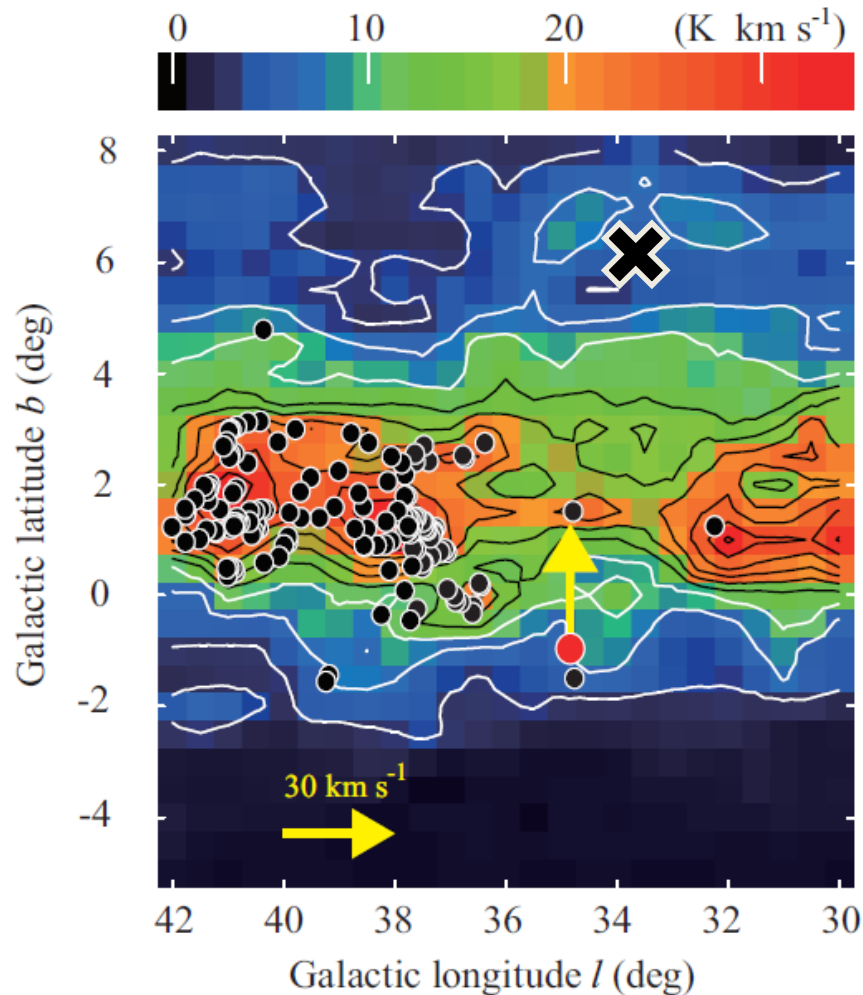
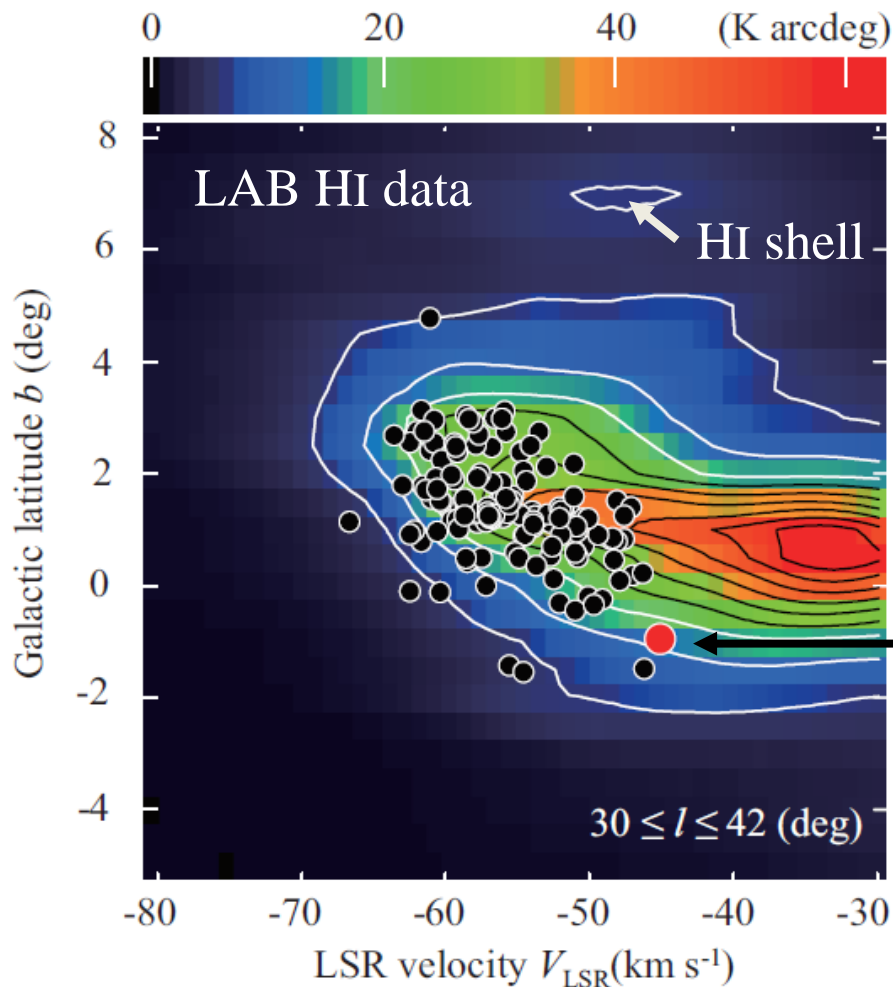
$$d_{V_{\text{LSR}}} = 18 +3/-2 \text{ kpc}$$

$$d_{\mu_l} = 20 +4/-3 \text{ kpc}$$

$$d_{3d} = 19 \pm 2 \text{ kpc (Preliminary)}$$

Pitch angle of the arm:  $11 \pm 2$  deg  
 (cf. Perseus arm =  $10 \pm 1$  deg)

# Discussions with the Preliminary results



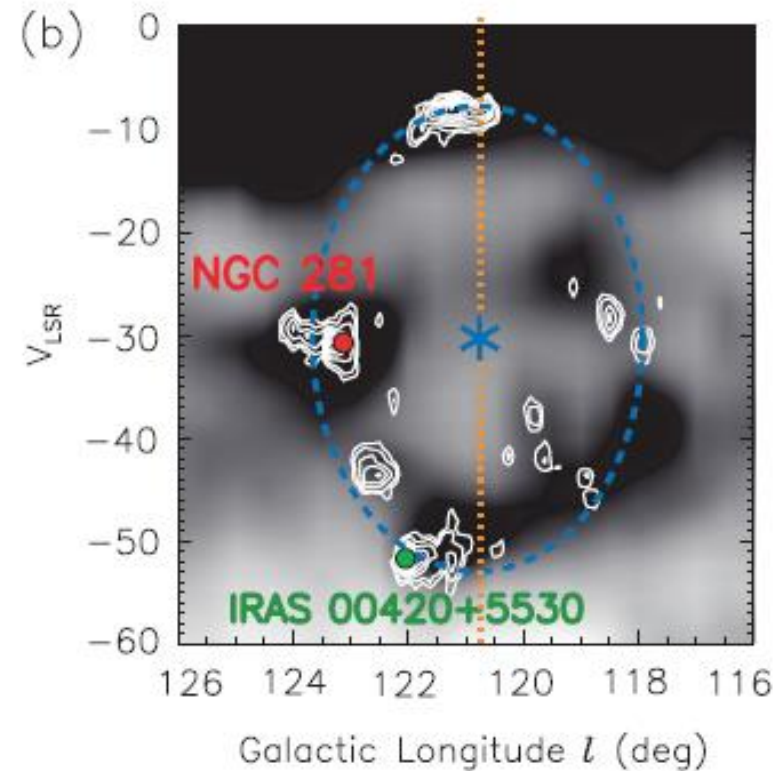
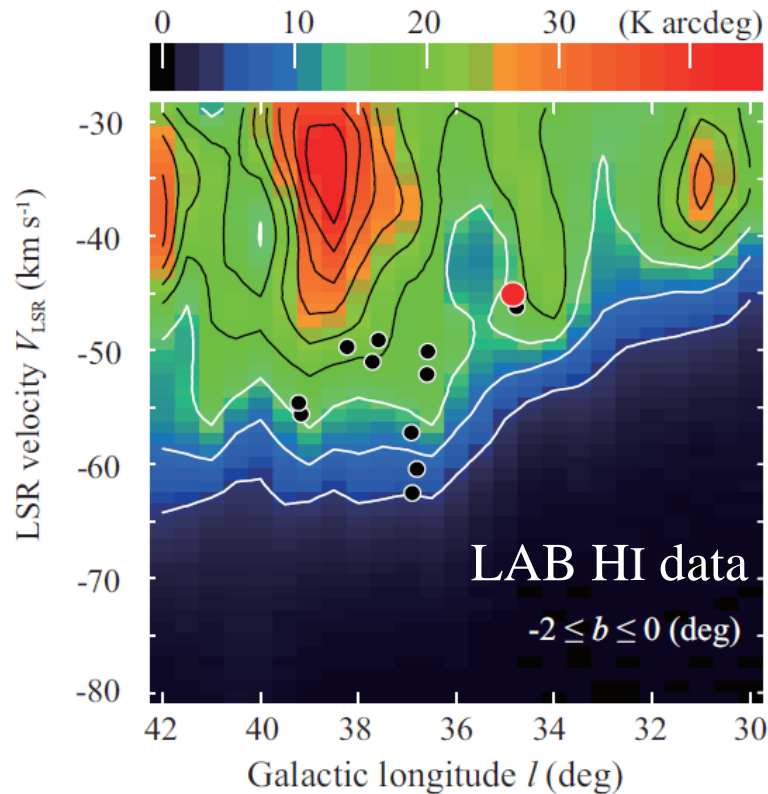
Galactic disk is warped toward positive  $b$   
 G034.84-00.95 position:  $z \sim -316 \pm 32$  pc

● CO (J=1-0) cloud (Dame+11; Sun+17)

✘ HI shell GS 033+06-49 (Heiles 1979)  
 $\mu_b + W_{\text{sun}} = \text{[redacted]} \text{ km s}^{-1}$   
 (insignificant at this moment)



# Discussions with the Preliminary results

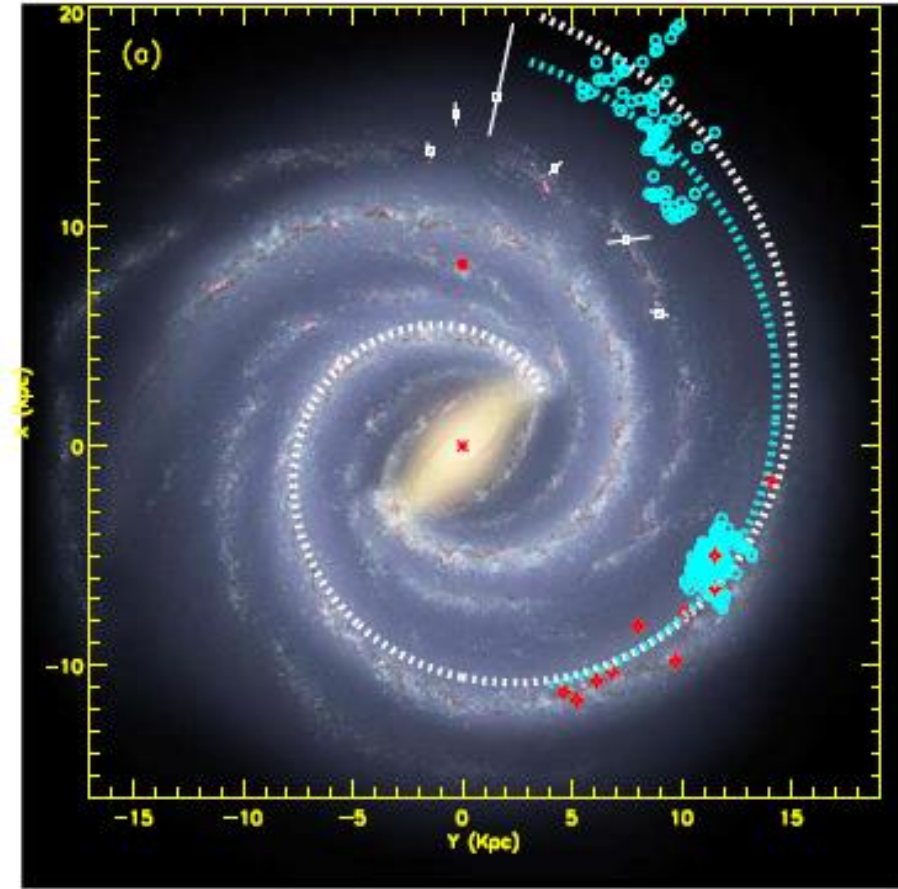


- G034 is associated with an edge of HI ring (left figure)
- $\Delta l \sim 400$  pc @  $d = 19$  kpc
- $\Delta V_{\text{LSR}} \sim 10$  km s<sup>-1</sup>
- $(\Delta l / 2) / (\Delta V_{\text{LSR}} / 2) \sim 40$  Myr

- NGC 281 and IRAS 00420 are associated with an edge of HI ring (right figure; Sato et al. 2008)
- $\Delta l \sim 270$  pc @  $d = 2.8$  kpc
- $\Delta V_{\text{LSR}} \sim 44$  km s<sup>-1</sup>
- $(\Delta l / 2) / (\Delta V_{\text{LSR}} / 2) \sim 6$  Myr

## Future

- ~281 Extreme Outer Galaxy CO clouds (Dame & Thaddeus 2011; Sun et al. 2015, 2017)
- ~5 masers detected toward brighter and massive EOG clouds, Sun et al. 2018 (Detection rate ~3%)
- EAVN 2021A CfP deadline:  
**2020/November/2<sup>nd</sup>**



Y. Sun, Y. Su, S.-B. Zhang, et al. 2017, *AplS*, 230, 17

# EOG astrometry by EAVN